

Henry J. G. L. M. Lamers *and*
Joseph P. Cassinelli

Introduction to STELLAR WINDS



Introduction To Stellar Winds

Chao Zhang



Introduction To Stellar Winds:

Introduction to Stellar Winds Henny J. G. L. M. Lamers, Joseph P. Cassinelli, 1999-06-17 The first comprehensive introduction to the observations and theories of stellar winds a long awaited graduate textbook written by two founders of the field *Hydrodynamics and Stellar Winds* Walter J. Maciel, 2014-02-11 Introduces hydrodynamics to undergraduate students in physics and astrophysics Stellar winds are a common phenomenon in the life of stars from the dwarfs like the Sun to the red giants and hot supergiants constituting one of the basic aspects of modern astrophysics Stellar winds are a hydrodynamic phenomenon in which circumstellar gases expand towards the interstellar medium This book presents an elementary introduction to the fundamentals of hydrodynamics with an application to the study of stellar winds The principles of hydrodynamics have many other applications so that the book can be used as an introduction to hydrodynamics for students of physics astrophysics and other related areas *Magnetohydrodynamics* J. P. Goedbloed, Hans

Goedbloed, Rony Keppens, Stefaan Poedts, 2019-01-31 An introduction to magnetohydrodynamics combining theory with advanced topics including the applications of plasma physics to thermonuclear fusion and plasma astrophysics **Cyclical Variability in Stellar Winds** Lex Kaper, Alexander W. Fullerton, 2013-06-05 It is well known that stellar winds are variable and the fluctuations are often cyclical in nature This property seems to be shared by the winds of cool and hot stars even though their outflows are driven by fundamentally different physical mechanisms Since very similar models have been proposed to explain the cyclical wind variations observed in a wide variety of stars the time was ripe for astrophysicists from many different sub disciplines to present the state of the art in a concise form The proceedings will provide a useful up to date overview of the observations interpretation and modelling of the time dependent mass outflows from all sorts of stars

Introduction to Stellar Structure Walter J. Maciel, 2015-09-16 This book provides a comprehensive overview of stellar structure evolution and basic stellar properties It includes integrated problems within the chapters with worked solutions In the first part of this book the author presents the basic properties of the stellar interior and describes them thoroughly along with deriving the main stellar structure equations of temperature density pressure and luminosity among others The process and application of solving these equations is explained as well as linking these results with actual observations The second part of the text describes what happens to a star over time and how to determine this by solving the same equations at different points during a star s lifetime The fate of various stars is quite different depending on their masses and this is described in the final parts of the book This text can be used for an upper level undergraduate course or an introductory graduate course on stellar physics *Encyclopedia of Astrobiology* Muriel Gargaud, Ricardo Amils, 2011-05-26 Astrobiology is a remarkably interdisciplinary field This reference serves as a key to understanding technical terms from the different subfields of astrobiology including astronomy biology chemistry the geosciences and the space sciences **Host Stars and their Effects on Exoplanet Atmospheres** Jeffrey Linsky, 2019-05-23 Like planets in our solar system exoplanets form

evolve and interact with their host stars in many ways As exoplanets acquire material and grow to the final size their atmospheres are subjected to intense UV and X radiation and high energy particle bombardment from the young host star Whether a planet can retain its atmosphere and the conditions for significant mass loss both depend upon the strength of the host star s high energy radiation and wind the distance of the exoplanet from its host star the gravitational potential of the exoplanet and the initial chemical composition of the exoplanet atmosphere This introductory overview describes the physical processes responsible for the emission of radiation and acceleration of winds of host stars that together control the environment of an exoplanet focusing on topics that are critically important for understanding exoplanetary atmospheres but are usually not posed from the perspective of host stars Accordingly both host stars and exoplanets are not studied in isolation but are treated as integrated systems Stellar magnetic fields which are the energy source for activity phenomena including high energy radiation and winds play a critical role in determining whether exoplanets are habitable This text is primarily for researchers and graduate students who are studying exoplanet atmospheres and habitability but who may not have a background in the physics and phenomenology of host stars that provide the environment in which exoplanets evolve It provides a comprehensive overview of this broad topic rather than going deeply into many technical aspects but includes a large list of references to guide those interested in pursuing these questions Nonspecialists with a scientific background should also find this text a valuable resource for understanding the critical issues of contemporary exoplanet research

Encyclopedia of Atmospheric Sciences Gerald R. North, John A. Pyle, Fuqing Zhang, 2014-09-14 Encyclopedia of Atmospheric Sciences Second Edition Six Volume Set is an authoritative resource covering all aspects of atmospheric sciences including both theory and applications With more than 320 articles and 1 600 figures and photographs this revised version of the award winning first edition offers comprehensive coverage of this important field The six volumes in this set contain broad ranging articles on topics such as atmospheric chemistry biogeochemical cycles boundary layers clouds general circulation global change mesoscale meteorology ozone radar satellite remote sensing and weather prediction The Encyclopedia is an ideal resource for academia government and industry in the fields of atmospheric ocean and environmental sciences It is written at a level that allows undergraduate students to understand the material while providing active researchers with the latest information in the field Covers all aspects of atmospheric sciences including both theory and applications Presents more than 320 articles and more than 1 600 figures and photographs Broad ranging articles include topics such as atmospheric chemistry biogeochemical cycles boundary layers clouds general circulation global change mesoscale meteorology ozone radar satellite remote sensing and weather prediction An ideal resource for academia government and industry in the fields of atmospheric ocean and environmental sciences *Models of Interacting Stellar Winds* Francis Patrick Wilkin, 1997 **The Astrophysics of Emission-Line Stars** Tomokazu Kogure, Kam-Ching Leung, 2010-05-05 Many types of stars show conspicuous emission lines in their optical spectra These stars are broadly

referred to as emission line stars Emission line stars are attractive to many people because of their spectacular phenomena and their amazing varieties and variability This book offers general information on emission line stars starting from a brief introduction to stellar astrophysics and then moving to a broad overview of emission line stars including early and late type stars as well as pre main sequence stars Detailed references and index are prepared for further reading Tomokazu Kogure s email address tkogure pa2 so net ne jp

The Observatory ,1923 A review of astronomy varies *Angular Momentum and Mass Loss for Hot Stars* L.A. Willson,R. Stalio,2012-12-06 Fundamental unsolved problems of stellar astrophysics include the effects of angular momentum on stellar structure and evolution the nature and efficiency of the processes by which angular momentum is redistributed within and lost from stars and the role that stellar rotation plays in enhancing or driving stellar mass loss There appears to be a qualitative change in the nature and efficiency of these mechanisms near spectral type FO hotter more massive stars typically retain more angular momentum at least until they reach the main sequence while cooler stars typically spin down quickly For the hotter stars recent work suggests a strong link between the type of pulsation behavior the mass loss rates and the rotation velocity If the same mechanisms are able to drive mass loss from the main sequence A stars as has recently been proposed then the current interpretations of a number of observations will be drastically affected e g the ages of clusters may be incorrect by up to a factor of two and the surface abundances of isotopes of He Li and Be may no longer give constraints on cosmological nucleosynthesis There are also effects on the evolution of the abundances of elements in the interstellar medium and on the general evolution of populations of stars Thus the questions of the mechanisms of angular momentum and mass loss of stars more massive than the sun is important not only for stellar studies but for the foundations of much of modern astrophysics

The Physics of the Interstellar Medium J E Dyson,D A Williams,1997-01-01 The book leads the advanced undergraduate through the wide range of disciplines related to an understanding of the interstellar medium and is suitable for any student studying either physics or astrophysics The study of the interstellar medium incorporates a large range of physical More processes on both large and small scales all of which are covered in this text Together with the inclusion of simple models and problems at the end of each chapter this text provides a comprehensive overview and grounding in the study of the interstellar medium

Literature 1982, Part 2 Siegfried Böhme,Professor Dr. Walter Fricke,Herbert Hefele,Inge Heinrich,Wilfried Hofmann,Dietlinde Krahn,Vladimir R. Matas,Dr. Lutz D. Schmadel,Gert Zech,2013-11-09 Astronomy and Astrophysics Abstracts which has appeared in semi annual volumes since 1969 is devoted to the recording summarizing and indexing of astronomical publications throughout the world It is prepared under the auspices of the International Astronomical Union according to a resolution adopted at the 14th General Assembly in 1970 Astronomy and Astrophysics Abstracts aims to present a comprehensive documenta tion of literature in all fields of astronomy and astrophysics Every effort will be made to ensure that the average time interval between the date of receipt of the original literature and publication of the abstracts will not exceed eight months This time interval is near to

that achieved by monthly abstracting journals compared to which our system of accumulating abstracts for about six months offers the advantage of greater convenience for the user Volume 32 contains literature published in 1982 and received before February 11 1983 some older literature which was received late and which is not recorded in earlier volumes is also included We acknowledge with thanks contributions to this volume by Dr J Bou a Prague who surveyed journals and publications in Czech and supplied us with abstracts in English

Mass Outflows from Stars and Galactic Nuclei Luciana Bianchi, Roberto Gilmozzi, 1988-04-30 The objective of this workshop was to put together observational and theoretical works on outflows from different kinds of astrophysical objects occurring on different scales and at various evolutionary phases and to discuss the impact of observations from future space missions For the stars we thought to follow throughout the evolution the relevance rates and dynamical modes of the mass loss phenomenon e g to explain how and when massive stars lose most of their initial mass to end up with typical WD masses The observations of the solar wind were included for being a unique case where the origin and propagation of the outflow can be resolved We thought that the comparison with similar phenomena occurring in galactic outflows would be fruitful as demonstrated by recent works on galactic winds and jets The interest of having this workshop in Torino came because there are groups in this area at the Astronomical Observatory and at the Institute of Physics of the University involved in the theoretical and observational studies of outflows from astrophysical objects The members of the Scientific Organizing Committee were V Castellani C Cesarski P Conti A Ferrari A Gabriel M Grewing Y Kondo H Lamers V Manno M Rees and R Schilizzi The Local Organizing Committee was L Bianchi G Massone and E Antonucci During the workshop the following topics were treated the solar wind the mass loss from cool stars and from hot stars

The Century of Space Science J.A. Bleeker, Johannes Geiss, M. Huber, 2012-12-06 One of the most attractive features of the young discipline of Space Science is that many of the original pioneers and key players involved are still available to describe their field Hence at this point in history we are in a unique position to gain first hand insight into the field and its development To this end *The Century of Space Science* a scholarly authoritative reference book presents a chapter by chapter retrospective of space science as studied in the 20th century The level is academic and focuses on key discoveries how these were arrived at their scientific consequences and how these discoveries advanced the thoughts of the key players involved With over 90 world class contributors such as James Van Allen Cornelis de Jager Eugene Parker Reimar Lüst and Ernst Stuhlinger and with a Foreword by Lodewijk Woltjer past ESO Director General this book will be immensely useful to readers in the fields of space science astronomy and the history of science Both academic institutions and researchers will find that this major reference work makes an invaluable addition to their collection

Pulsation, Rotation and Mass Loss in Early-Type Stars Luis A. Balona, Huib F. Henrichs, Jean Michel Le Contel, 2012-12-06 In this Symposium researchers specializing in pulsation rotation magnetic fields and stellar winds are brought together for the first time in order to broaden our understanding of O and B stars Thanks to advances in digital spectroscopy new types of pulsating B

stars have been discovered The pulsations can be understood in terms of the recent revision of metal opacities but the effects of rapid rotation and magnetic fields need further study Observations in the UV and X ray regions demonstrate that many B and Be stars show other activity besides pulsation which is not yet understood The reason for the enhanced mass loss in Be stars is a question which dominates the Symposium and which remains unanswered although it is surely to be found in activity at or near the photosphere coupled with rotation It is shown that the geometry of the circumstellar envelopes around Be stars is indeed a flattened disk as they can now be optically resolved The variability of radiatively driven winds from O and B stars are likely related to the rotation of the star This underlines the central theme of the book that the various phenomena seen in these stars cannot be studied in isolation Star-Disk Interaction in Young Stars (IAU S243) International Astronomical Union. Symposium,2007 Reviews latest star disk interaction region in young stars for graduate students and researchers Physics and Chemistry of the Interstellar Medium Sun Kwok,2007-01-30 This book goes beyond a phenomenological study to present a detailed quantitative treatment of the dynamic interactions between stars and interstellar matter Emphasizing a practical understanding of these processes the text is interlaced with mathematical derivations that are understandable by anyone with an undergraduate background in Physics **Evolution of Interstellar Matter and Dynamics of Galaxies** J. Palous,B. W. Burton,P. O. Lindblad,1992-07-02 Proceedings of the international workshop on the Evolution of Interstellar Matter and Dynamics of Galaxies

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